Numbers You Should Know Part 1

A semi-quantitative profile of the profession

ASTR 8500

Spring 2022

Robert W. O'Connell

Sources?

- American Astronomical Society
- American Institute of Physics
- NSF, NASA, other agencies
- UVa documents
- Literature and general media sources
- ROMEs = rough order magnitude estimates

ASTR 8500 (O'Connell, Spring 2022)

PRESENTATIONS AND RESOURCES

A. Course Assignments

B. Numbers You Should Know (O'Connell)

<u>NSF Outlook</u> (AAS Town Hall, Jan 2020) <u>NASA Outlook</u> (AAS Town Hall, Jan 2020)

NSF Outlook (AAS Town Hall, Jan 2022) NASA Outlook (AAS Town Hall, Jan 2022)

"Facts of Life for New Teachers in the Astronomy Nonmajors Curriculum" (O'Connell, AstEdRev, 6, 1, 2007) "Production Rate and Employment of PhD Astronomers" (Metcalf, PASP, 120, 229, 2008) "A Closer Look at Astronomy Faculty" (Ivie, AIP, 2009) BA Degree Gender Gap By Field, 1970-2010 (Olson graphic, 2013) "Astronomy Enrollments and Degrees" (Mulvey & Nicolson, AIP, 2014) "Longitudinal Study of Astronomy Graduate Students" (Ivie, AIP, 2014) "Doctorate Recipients from US Universities" (NSF, 2015) "Astronomy Degree Recipients Initial Employment" (Pold, AIP, 2015) "Women's and Men's Career Choices in Astronomy and Astrophysics" (Ivie et al., PRPER, 12, 020109, 2016) "A Survey of U.S. Astronomers" (Spuck, PhD Thesis, WVU, 2017) "Degree Plus One, Employment & Salaries" (Mulvey & Pold, AIP, 2019) "Long Term Trends in the Astronomical Workforce" (Momcheva, Astro 2020 White Paper) "Gender & the Career Outcomes of PhD Astronomers in the US" (Perley, PASP, 131:114502, 2019)

C. Navigating the Early Career Job Market (Meyer and McGuire)

AAS Job Register Astronomy Rumor Mill (astrobetter)

"The Professor Is In: The Essential Guide to Turning Your Ph.D. Into a Job" (Kelsky)

D. Consolidated Faculty Advice for Graduate Students

"Tips for Success in Observational Astronomy" (O'Connell)

A National Perspective on Astronomy

We're #3 among STEMM fields in media impact:

#1 Health & Medicine#2 Environment#3 Astronomy & Space

We're #1 in impact per practitioner!

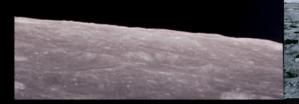
Total STEMM employment: 16 million

Total astronomy & space science employment: ~20-30 thousand

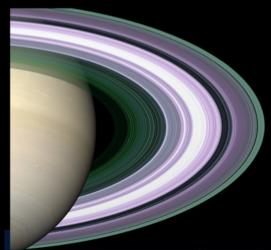


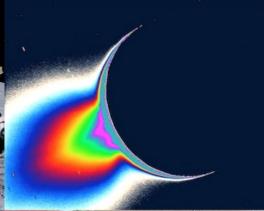
















SUN 9/8c MAR 9 COSMOS: A SPACETIME ODYSSEY Presented by FOX Sun 9/8c and National Geographic Mon 10/9c

Samsung GALAXY

Jeep





Blog Clips Live Event

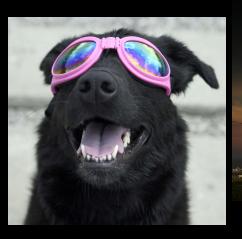


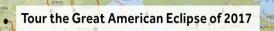
SUNDAY MARCH 9 ON 10 CHANNELS FOX ANTIONAL MATCHO MATCHO FOX LIFE FX 5% FXM (2) (2)

🛛 🛉 🕈 🕈 🗾



Eclipse USA 2017





A free web app by





The New York Times

215 Million Americans Watched the Solar Eclipse, Study Finds



Todd Heisler/The New York Times

By Jonah Engel Bromwich

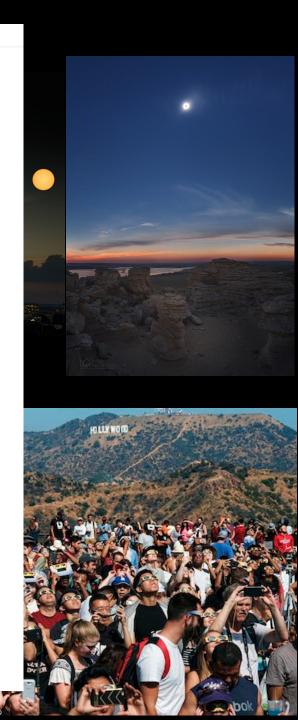
Sept. 27, 2017

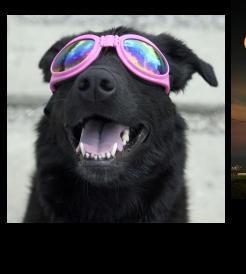


We hear it all the time: Americans are <u>more divided than ever</u>, or at least since the Civil War.

But <u>the solar eclipse on Aug. 21</u> brought the United States together in greater numbers than most any national event in recent memory, according to a study released Tuesday by the University of Michigan. It estimated that 88 percent of American adults about 215 million people — watched the solar eclipse, either in person or electronically.

That's nearly twice the number of people that watched the Super Bowl last year. It's <u>almost 30 percent more Americans</u> than participated in the presidential election last year.



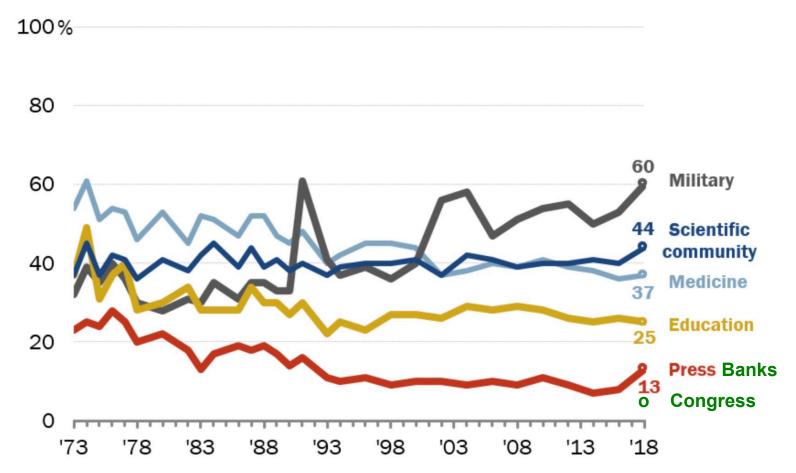




"Scientist" is the #4 most prestigious profession (Harris Poll, 2014)

Confidence in leaders of the military has gone up; confidence in some other institutions is declining

% of U.S. adults who say they have a great deal of confidence in the people running each of these institutions



Note: Respondents who gave other responses or who did not give an answer are not shown. Source: General Social Surveys, NORC.

PEW RESEARCH CENTER (2018 data)

250,000

= Number of college students enrolled annually in elementary astronomy courses





\$425,000

= National budget for astronomy per astronomer

THE NATIONAL BUDGET FOR ASTRONOMY (2016)

NSF	NASA	DOE, DOD	Univ/Priv*	Total**	Number Astronomers***	\$\$/Astronomer
\$250M	\$2950M	~\$50M	~\$150M	\$3400M	~8000	\$425,000

^{*}Research support; excludes basic faculty salaries.

**The federal budget for astronomy is ~0.08% of the total federal budget of \$4.0T or \$10.09 per US citizen per year.

**AAS membership, 2016

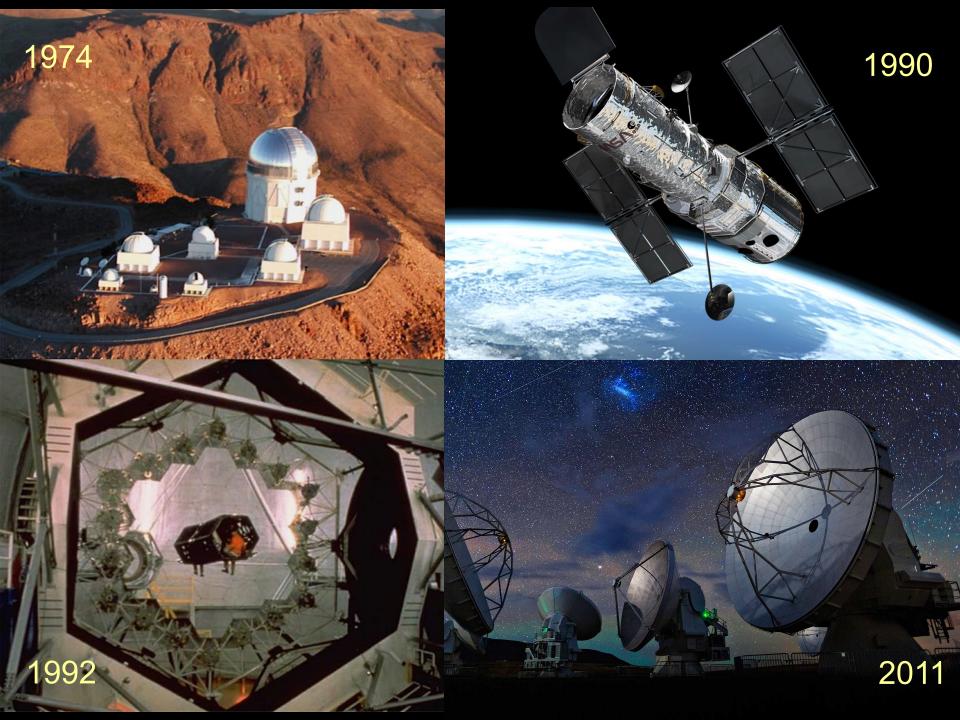




\$425,000

= National budget for astronomy per astronomer

...BUT: mostly in the form of <u>shared</u> observing facilities





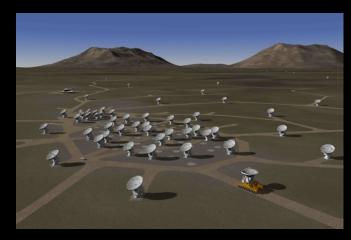


Astronomy is ~uniquely dependent on large, shared experimental facilities

Astronomical facilities have <u>long</u> productive lifetimes; a blessing & a curse



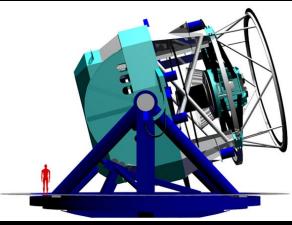
Promise of the Next Decade



ALMA



GMT



Rubin/LSST

JWST



Roman/WFIRST

The Job Market

WHEW!



344 possible single point failures

....all avoided

\$60M/year for GO programs



Kinds of Jobs for Astronomers

Postdoctoral

Short-term (1-3 yr) research positions (mostly directed)

Research Scientists

 Mostly semi-permanent. Large range, from support to independent researchers. Universities, observatories, government labs (e.g. NRAO, NOIRLab, GSFC, STScI, USNO). Independent contractors (e.g. JHU/APL, SWRI)

University Faculty

- Short term contractual and permanent (tenured)
 Research and teaching
- Non-astro-research Government
 - E.g. NASA, NSF, DOD, DOE, NOAA, etc.
- Non-astro-research Private Sector
 - "Beltway Bandits," high-end computing, aerospace,
 - sensors & optics, medical imaging, communications...

Kinds of \$\$\$

"Hard" money (reliable, long-term)

- Tenured faculty
- Civil servants
- Tenured & senior staff at national labs
- "Soft" money (term-limited, grants, contracts)
 - Postdocs
 - "Adjunct" faculty
 - Many "research scientists"
 - Federal contractors (e.g. SWRI)
 - Other private sector

= Unemployment rate for astronomers ~ Transition rate

2%

Full employment

Astronomy Long-term PhD Employment Pattern Through the 1990's:

~1/3 Faculty
~1/3 Research Scientists
~1/3 Non Astronomy

(Perley 2019)

Astronomy Long-term PhD Employment Pattern Through the 2000's:

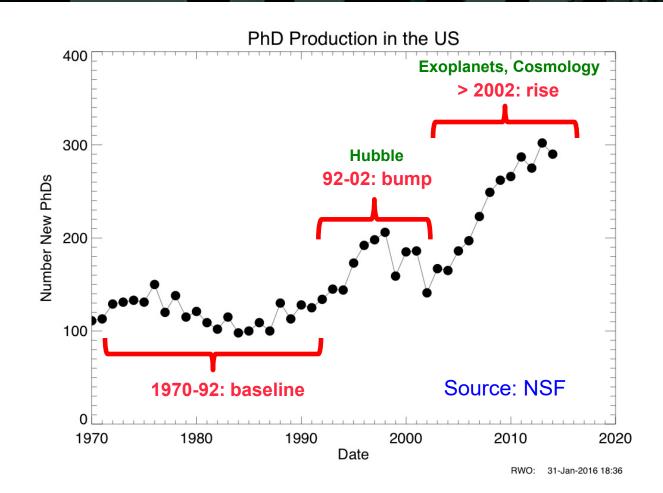
~1/3 Faculty
 ~1/3 Research Scientists
 ~1/3 Non Astronomy

(Perley 2019)

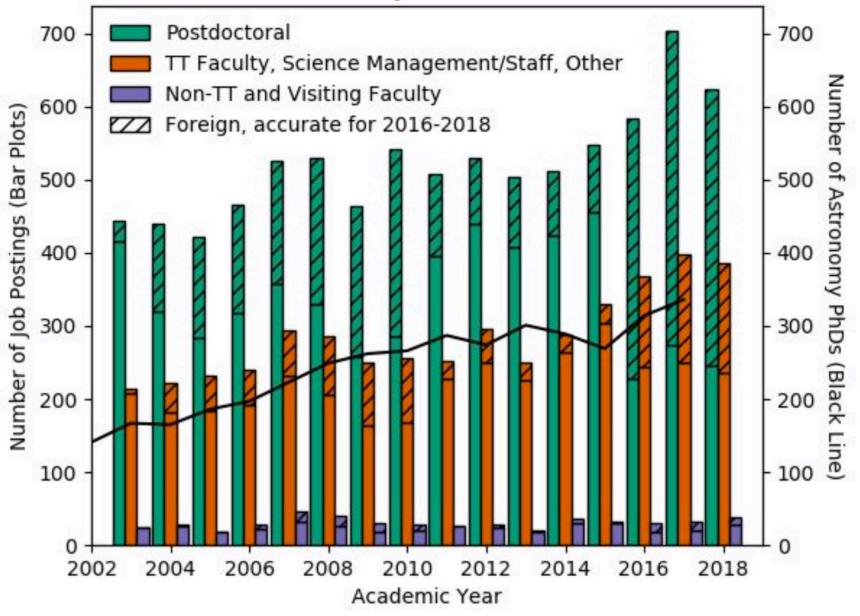
How many jobs? ~8500

Membership of AAS + NonmembersNon-grad degrees

Production of New Astronomy PhD's



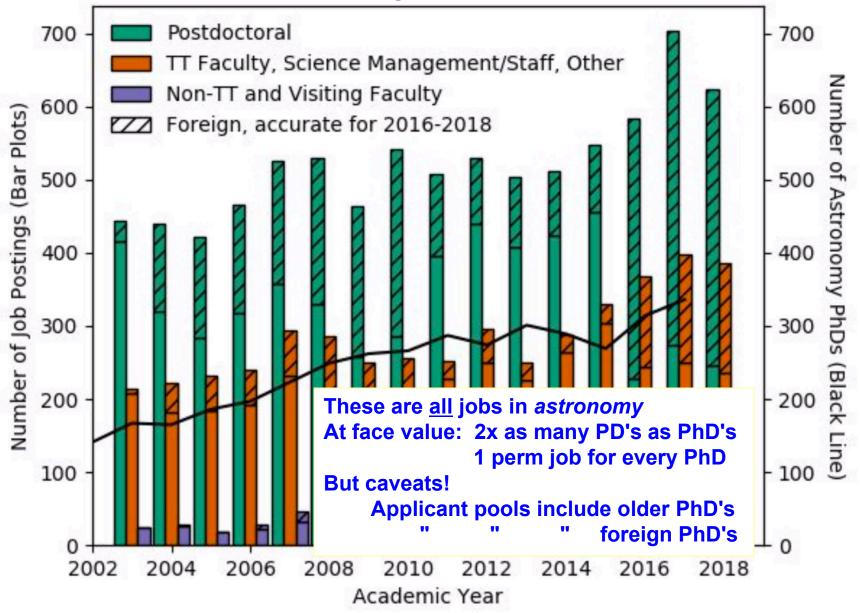
AAS Job Register Statistics



Black line: PhD's produced

(Kamenetzky, White Paper, 2019)

AAS Job Register Statistics



Black line: PhD's produced

(Kamenetzky, White Paper, 2019)

Employment Demographics Studies

Perley 2019

Publications of the Astronomical Society of the Pacific, 131:114502 (7pp), 2019 November © 2019. The Astronomical Society of the Pacific, All rights reserved. Printed in the U.S.A. https://doi.org/10.1088/1538-3873/ab0cc4



Gender and the Career Outcomes of Ph.D. Astronomers in the United States

Daniel A. Perley 😳

Astrophysics Research Institute, Liverpool John Moores University, IC2, Liverpool Science Park, 146 Brownlow Hill, Liverpool L3 5RF, UK; d.a. perley@ljmu.ac.uk Received 2019 February 10; accepted 2019 February 25; published 2019 September 24

Abstract

We analyze the postdoctoral career tracks of a nearly complete sample of astronomers from 28 United States graduate astronomy and astrophysics programs spanning 13 graduating years (N = 1063). A majority of both men and women (65% and 66%, respectively) find long-term employment in astronomy or closely related academic disciplines. We find no significant difference in the rates at which men and women are hired into these jobs following their Ph.D.s or in the rates at which they leave the field. Applying a two-outcome survival analysis model to the entire data set, we measure a relative academic hiring probability ratio for women versus men at a common year -post-Ph.D. of $H_{F/M} = 1.08^{+0.20}_{-0.17}$ and a leaving probability ratio of $L_{F/M} = 1.03^{+0.31}_{-0.24}$ (95% CI). These are both consistent with equal outcomes for both genders ($H_{F/M} = L_{F/M} = 1$) and rule out more than minor gender differences in hiring or in the decision to abandon an academic career. They suggest that despite discrimination and adversity, women scientists are successful at managing the transition between Ph.D., postdoctoral, and faculty/ staff positions.

Key words: sociology of astronomy

Online material: color figures, machine-readable table

Employment Demographics Studies

Perley 2019

Publications of the Astronomical Society of the Pacific, 131:114502 (7pp), 2019 November © 2019. The Astronomical Society of the Pacific. All rights reserved. Printed in the U.S.A. https://doi.org/10.1088/1538-3873/ab0cc4



Gender and the Career Outcomes of Ph.D. Astronomers in the United States

Daniel A. Perley 😳

Astrophysics Research Institute, Liverpool John Moores University, IC2, Liverpool Science Park, 146 Brownlow Hill, Liverpool L3 5RF, UK; d.a. perley@ljmu.ac.uk Received 2019 February 10; accepted 2019 February 25; published 2019 September 24

Followed actual post-PhD histories of over 1100 individual astronomers, <u>unlike</u> earlier studies using broad statistical measures or unreliable reporting (e.g. the "Rumor Mill" site).

Data for 2000-2012.



12-Year Post-PhD Statistics (Data 2000-2012)

Frac starting perm job

Frac leaving field

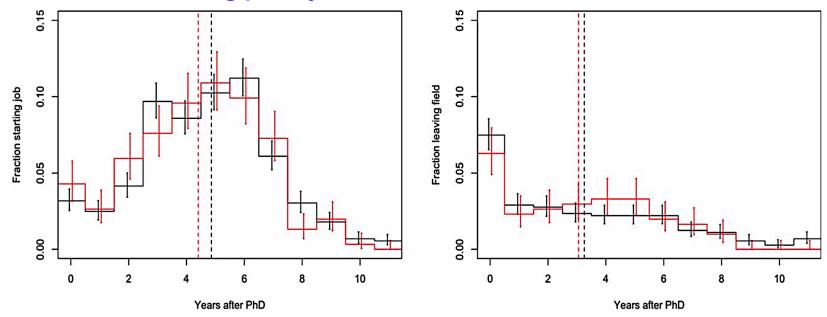
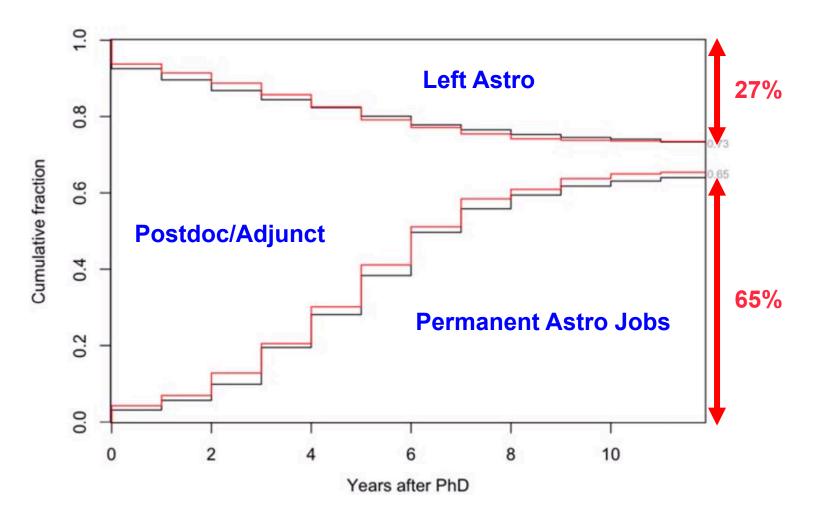


Figure 1. Histograms of recorded times (years after Ph.D.) at which Ph.D.s either: (left) progressed from term-limited to long-term or permanent positions within astronomy, or (right) left the field to pursue other employment. Histograms are normalized using total counts for each gender (regardless of outcome). Error bars show 67% binomial confidence intervals and dashed vertical lines show the means. Male astronomers are shown in black and female astronomers in red.



Perley 2019

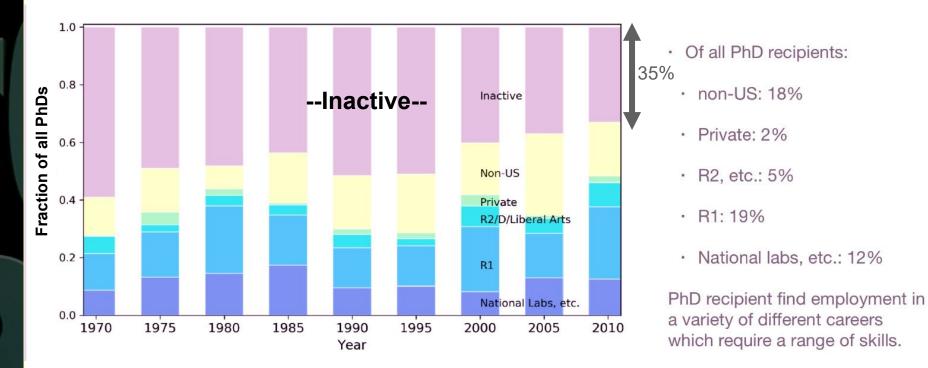
12-Year Cumulative Employment (Data 2000-2012)



Employment Demographics Studies

Momcheva 2019

"Active" = published in professional literature within 3 yrs



Perley (2019): "The number of astronomy PhD's is not greatly in excess of the number of careers available within the field." [Data for 2000-2012]

Kamenetzky (2019): "The overall number of potentially permanent positions...has slightly increased in the past decade [2010-2019] to~380 per year compared to ~270 ten years earlier, roughly keeping pace with the increase in new PhDs."

~ Permanent jobs in *astronomy*.

Perley (2019): "The number of astronomy PhD's is in excess of the number of care within the field." [Data for 2010]

Kamenetzky (2019): "The overal" manent posit" the past

CO

ally perincreased in all to ~380 per year years earlier, roughly the increase in new PhDs." Astronomy Long-term PhD Employment Pattern Through the 2000's:

~1/3 Faculty
 ~1/3 Research Scientists
 ~1/3 Non Astronomy

(Perley 2019)

Positions Held by UVa Astro PhD's (1967-2017; 130 Degrees)

Faculty Research scientists Non-Astro Outreach Secondary ed Postdocs

33% 40% 13% 2% 1% 1%

Numbers You Should Know Part 2

Professional effort

ASTR 8500

Spring 2022

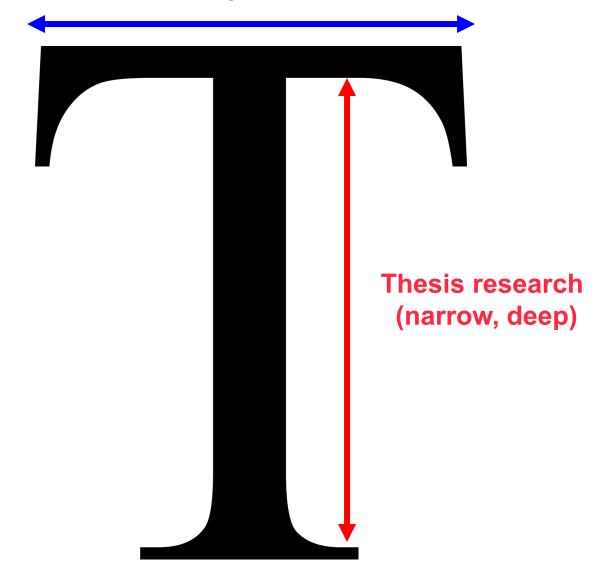
Robert W. O'Connell

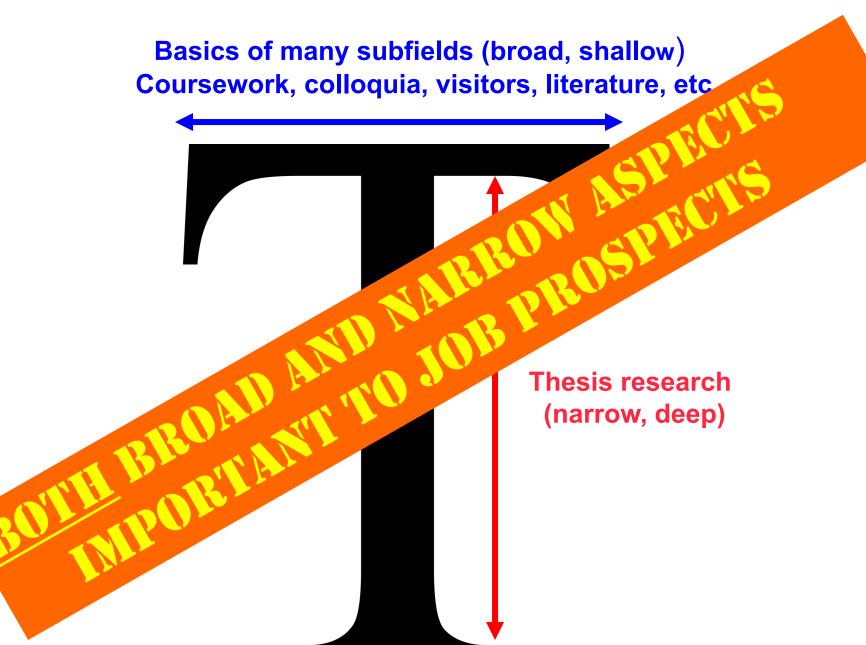
~95(?)% 95(?)% of astronomers love most of what they do

Life as a Graduate Student



Basics of many subfields (broad, shallow) Coursework, colloquia, visitors, literature, etc





10,000

= Number of HOURS of close engagement with a speciality before a person is ready to make important contributions

May involve actual *reconfiguration* of neural circuits(!)

10,000

= Number of HC with a speci make in

erson is ready to

actual reconfiguration of neural

Life as a Faculty Member



Responsibilities of a faculty member

13

a management and and the

Teaching

•

- Classroom teaching (mostly undergrad 90-95% nonmajors)
- Tutorial, small group instruction
- Course, curriculum, & resource development/management
- Student mentoring, advising, recommendations
- Outreach
- Research
 - Personal undirected*
 - Supervising grad student & postdoc research
 - Management: lab/group direction, obtaining & administering finances (grants)

Service/Administration

 Local department & university administration: operations, governance, policies, personnel evaluation (recruiting, promotions)

TATALAN DISTANCE PROPERTY AND THE DESIGNATION OF TH

- Refereeing publications, proposal reviews
- Disciplinary activities, planning, meetings, advocacy
- National agency policy, planning, review
- Consulting

Teaching

•

- Classroom teaching (mostly undergrad 90-95% nonmajors)
- Tutorial, small group instruction
- Course, curriculum, & resource development/management
- Student mentoring, advising, recommendations
- Outreach
- Research
 - Personal undirected*
 - Supervising grad student & postdoc research
 - Management: lab/group direction, obtaining & administering finances (grants)

Service/Administration

- Local department & university administration: operations, governance, policies, personnel evaluation (recruiting, promotions)
- Refereeing publications, proposal reviews
- Disciplinary activities, planning, meetings, advocacy
- National agency policy, planning, review
- Consulting

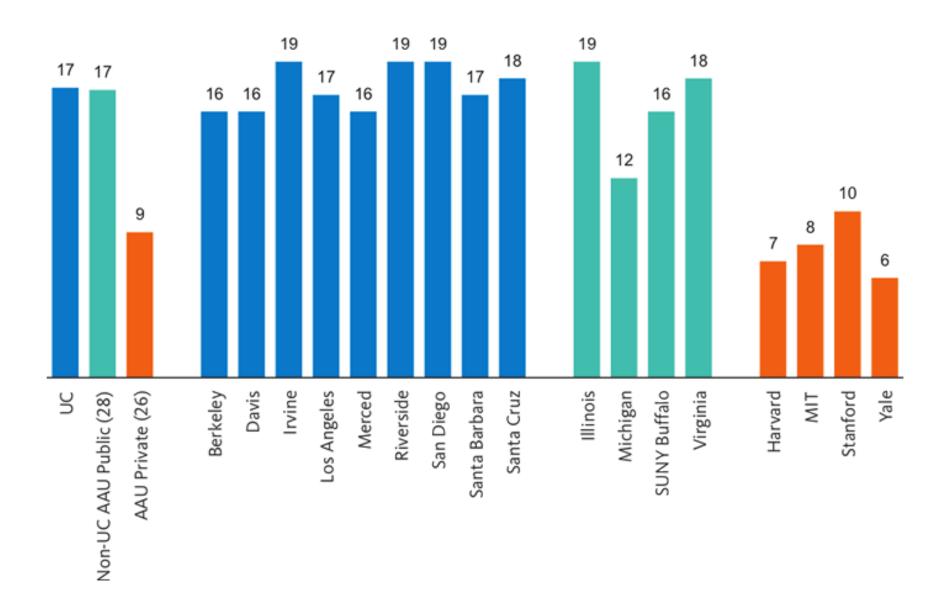
Almost no middle management

TRANSFORMENT CONTRACTOR AND INCOMENTS

~90-200

= Number of UG students the average professor must teach each year
= 10 x S/F (in 3-credit classes)

Student/Faculty Ratios



TEACHING LOADS: THE RELENTLESS ARITHMETIC

If S/F = 20 and if half your department faculty teaches "small" classes, with 25 students, then the other half of the faculty must teach, on average, 175 students per semester. I.e. **SEVEN TIMES MORE STUDENTS.**

= Career averaged ratio of <u>total</u> teaching time to <u>in-class</u> time

~7:1

Averaged effort for a 3-hour course: ~20 hours/wk

New burden for teachers! Increasing emphasis on ELECTRONICS time-lapse projector oracle asur <u>Bed</u> e-grai vime IOCX product-cycle

* One hour of course video takes 50-100 hours of prep

1:1 to 3:1 = Ratio of <u>real-time</u> rehearsal to delivery time for a wellprepped talk

- An important class lecture
- A job talk
- A review talk
- A news conference
- etc

D Ideas worth spreading

- An important class lecture
- A job talk
- A review talk
- A news conference
- etc

Diden route man route

- A review talk
- A news conference
- etc

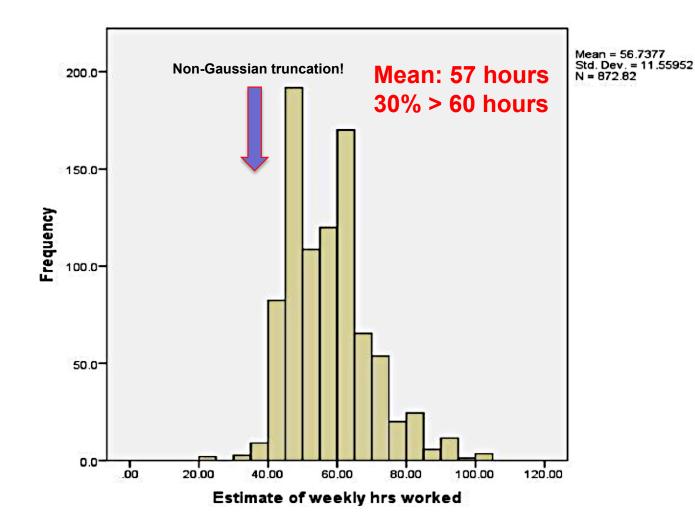




= Number of hours per week professors *claim* to work

UVa Faculty Senate Survey (2012)

Figure VII-1: Frequency Distribution of Hours Worked Per Week, Full Time Faculty Only.

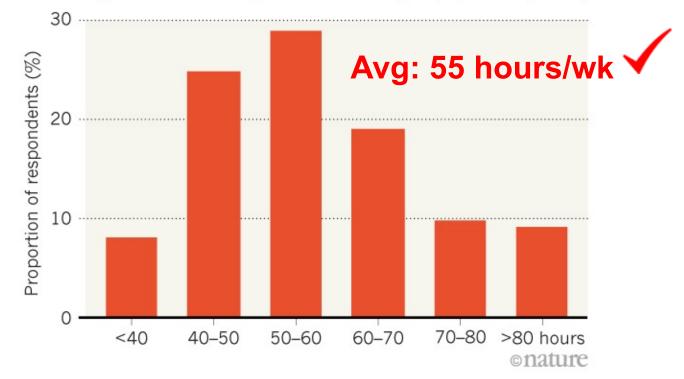


LONG HOURS

Some 38% of *Nature*'s readers say they work more than 60 hours a week.

Poll question:

How many hours a week do you work on average? (12,869 responses)









Teaching

•

- Classroom teaching (mostly undergrad)
- Tutorial, small group instruction
- Course, curriculum, & resource development/management
- Student mentoring, advising, recommendations
- Outreach
- Research
 - Personal undirected
 - Supervising grad student & postdoc research
 - Management: lab/group direction, obtaining & administering finances (grants)

Service/Administration

 Local department & university administration: operations, governance, policies, personnel evaluation (recruiting, promotions)

- Refereeing publications, proposal reviews
- Disciplinary activities, planning, meetings, advocacy
- National agency policy, planning, review
- Consulting

Teaching

•

- Classroom teaching (mostly undergrad)
- Tutorial, small group instruction
- Course, curriculum, & resource development/manage
- Student mentoring, advising, recommendations
- Outreach
- Research
 - Personal undirected
 - Supervising grad stude
 - Management: la
 - Service/Adr
 - Local pers
 - Refereen
 - Disciplinar

astration: operations, governance, policies, g, promotions)

ministering finances (grants)

TITLET DESCRIPTION ADDRESS 111

- oposal reviews
- , planning, meetings, advocacy
- National age _y policy, planning, review
- Consulting

no man man man man

Teaching

•

- Classroom teaching (mostly undergrad)
- Tutorial, small group instruction
- Course, curriculum, & resource development/manage
- Student mentoring, advising, recommendations
- Outreach

Research •

- Personal -undirected
- Supervising grad stude
- Management: lak
- Service/Adr

Refe

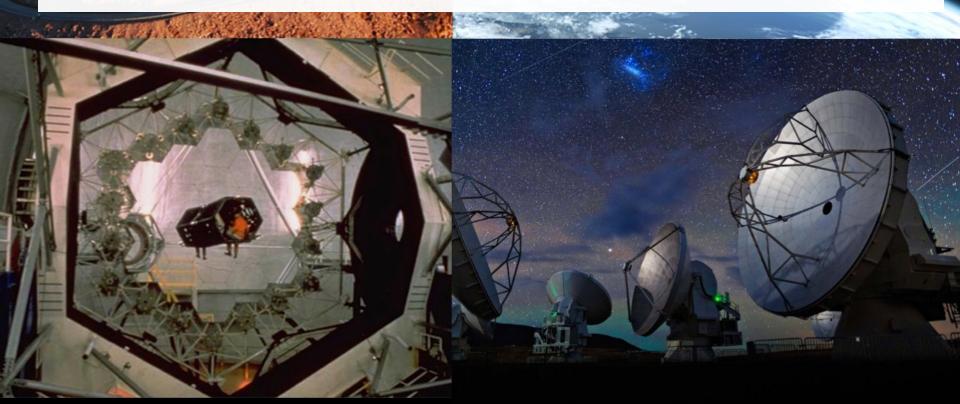
mnistering finances (grants)

TELEVISION DESCRIPTION OF TAXABLE INC.

E CANAL YOUR TIME stration: operations, governance, policies,

- Discipl planning, meetings, advocacy
- National y policy, planning, review
- Consulting tett (tetettet greatet 32.

Life as a Research Scientist



Job Profile of a Research Scientist

Research Support

- Observer support & training
- Telescope time allocation
- .-. Software design, development, oversight.
- Data analysis pipelines, data archives, quality assurance
- Instrumentation development
- Documentation
- Facility upgrade projects
- Policy formulation
- Personnel administration
- Personal Research
 - Allocation usually specified; typically 15-50% but wide variation
 - Grant support provides buy-outs of service time
 - General Service
 - Refereeing publications, proposal reviews.
 - Disciplinary activities, planning, meetings, advocacy
 - National agency policy, planning, review
 - Consulting

Job Profile of a Research Scientist

Main occupational hazard of a research scientist?

Meetings

Astronomy Long-term PhD Employment Pattern Through the 2000's:

~1/3 Faculty ~1/3 Research Scientists ~1/3 Non Astronomy

(Perley 2019)

65%

"Non-Dedicated Astronomy"

Jobs drawing on general training in high-tech field Examples:

Space science/applications (govt, contractors, commercial) High-end computing (databases, AI) Computational biology (genomics, neurology) Communications (radio, microwave, fiber/laser) Instrumentation (sensors, imaging, optics) Medical imaging

Be alert for opportunities to develop transferable knowledge and skills

= Largest number of authors on an astronomical paper

3611

The Rise of Group Science

ADS Statistics on published Ast/Ap papers

 # Authors
 1975
 2016

 1
 40%
 7%

 >2
 26%
 78%

 >5
 3%
 39%

 Max # to date
 54
 1187

3611
 Largest number of authors on an astronomical paper.

Abbott et al., "Multi-Messenger Observations of a Binary Neutron Star Merger," ApJL, 848, L12, 2017 Abbott et all ! THE ASTROPHYSICAL JOURNAL LETTERS, 848:L12 (59pp), 2017 October 20
© 2017. The American Astronomical Society. All rights reserved.
OPEN ACCESS



Multi-messenger Observations of a Binary Neutron Star Merger*

LIGO Scientific Collaboration and Virgo Collaboration, Fermi GBM, INTEGRAL, IceCube Collaboration, AstroSat Cadmium Zinc Telluride Imager Team, IPN Collaboration, The Insight-HXMT Collaboration, ANTARES Collaboration, The Swift Collaboration, AGILE Team, The 1M2H Team, The Dark Energy Camera GW-EM Collaboration and the DES Collaboration, The DLT40 Collaboration, GRAWITA: GRAvitational Wave Inaf TeAm, The Fermi Large Area Telescope Collaboration, ATCA: Australia Telescope Compact Array, ASKAP: Australian SKA Pathfinder, Las Cumbres Observatory Group, OzGrav, DWF (Deeper, Wider, Faster Program), AST3, and CAASTRO Collaborations, The VINROUGE Collaboration, MASTER Collaboration, J-GEM, GROWTH, JAGWAR, Caltech-NRAO, TTU-NRAO, and NuSTAR Collaborations, Pan-STARRS, The MAXI Team, TZAC Consortium, KU Collaboration, Nordic Optical Telescope, ePESSTO, GROND, Texas Tech University, SALT Group, TOROS: Transient Robotic Observatory of the South Collaboration, The BOOTES Collaboration, LWA: Murchison Widefield Array, The CALET Collaboration, TKI-GW Follow-up Collaboration, ALMA Collaboration, LOFAR Collaboration, LWA: Long Wavelength Array, HAWC Collaboration, The Pierre Auger Collaboration, ALMA Collaboration, LOFAR Collaboration, UWA: Long Wavelength Array, HAWC Collaboration, The Pierre Auger Collaboration, ALMA Collaboration, Euro VLBI Team, Pi of the Sky Collaboration fre Chandra Team at McGill University, DFN: Desert Fireball Network, ATLAS, High Time Resolution Universe Survey, RIMAS and RATIR, and SKA South Africa/MeerKAT (See the end matter for the full list of authors.)

Received 2017 October 3; revised 2017 October 6; accepted 2017 October 6; published 2017 October 16

Abstract

On 2017 August 17 a binary neutron star coalescence candidate (later designated GW170817) with merger time 12:41:04 UTC was observed through gravitational waves by the Advanced LIGO and Advanced Virgo detectors. The Fermi Gamma-ray Burst Monitor independently detected a gamma-ray burst (GRB 170817A) with a time delay of \sim 1.7 s with respect to the merger time. From the gravitational-wave signal, the source was initially localized to a sky region of 31 deg² at a luminosity distance of 40^{+8}_{-8} Mpc and with component masses consistent with neutron stars. The component masses were later measured to be in the range 0.86 to 2.26 M_{\odot} . An extensive observing campaign was launched across the electromagnetic spectrum leading to the discovery of a bright optical transient (SSS17a, now with the IAU identification of AT 2017gfo) in NGC 4993 (at ~40 Mpc) less than 11 hours after the merger by the One-Meter, Two Hemisphere (1M2H) team using the 1 m Swope Telescope. The optical transient was independently detected by multiple teams within an hour. Subsequent observations targeted the object and its environment. Early ultraviolet observations revealed a blue transient that faded within 48 hours. Optical and infrared observations showed a redward evolution over ~10 days. Following early non-detections, X-ray and radio emission were discovered at the transient's position ~ 9 and ~ 16 days, respectively, after the merger. Both the X-ray and radio emission likely arise from a physical process that is distinct from the one that generates the UV/optical/near-infrared emission. No ultra-high-energy gamma-rays and no neutrino candidates consistent with the source were found in follow-up searches. These observations support the hypothesis that GW170817 was produced by the merger of two neutron stars in NGC 4993 followed by a short gamma-ray burst (GRB 170817A) and a kilonova/macronova powered by the radioactive decay of r-process nuclei synthesized in the ejecta.

Key words: gravitational waves - stars: neutron

1. Introduction

Over 80 years ago Baade & Zwicky (1934) proposed the idea of neutron stars, and soon after, Oppenheimer & Volkoff (1939) carried out the first calculations of neutron star models. Neutron stars entered the realm of observational astronomy in the 1960s by providing a physical interpretation of X-ray emission from Scorpius X-1 (Giacconi et al. 1962; Shklovsky 1967) and of radio pulsars (Gold 1968; Hewish et al. 1968; Gold 1969).

The discovery of a radio pulsar in a double neutron star system by Hulse & Taylor (1975) led to a renewed interest in binary stars and compact-object astrophysics, including the

^{*} Any correspondence should be addressed to lvc.publications@ligo.org.



development of a scenario for the formation of double neutron stars and the first population studies (Flannery & van den Heuvel 1975; Massevitch et al. 1976; Clark 1979; Clark et al. 1979; Dewey & Cordes 1987; Lipunov et al. 1987; for reviews see Kalogera et al. 2007; Postnov & Yungelson 2014). The Hulse-Taylor pulsar provided the first firm evidence (Taylor & Weisberg 1982) of the existence of gravitational waves (Einstein 1916, 1918) and sparked a renaissance of observational tests of general relativity (Damour & Taylor 1991, 1992; Taylor et al. 1992; Wex 2014). Merging binary neutron stars (BNSs) were quickly recognized to be promising sources of detectable gravitational waves, making them a primary target for groundbased interferometric detectors (see Abadie et al. 2010 for an overview). This motivated the development of accurate models for the two-body, general-relativistic dynamics (Blanchet et al. 1995; Buonanno & Damour 1999; Pretorius 2005; Baker et al. 2006; Campanelli et al. 2006; Blanchet 2014) that are critical for detecting and interpreting gravitational waves (Abbott et al. 2016c, 2016d, 2016e, 2017a, 2017c, 2017d).

GROUPS, not people



B. P. Abbott¹, R. Abbott¹, T. D. Abbott², F. Acemese^{3,4}, K. Ackley^{5,6}, C. Adams⁷, T. Adams⁸, P. Addesso⁹, R. X. Adhikari¹, V. B. Adya¹⁰, C. Affeldt¹⁰, M. Afrough¹¹, B. Agarwal¹², M. Agathos¹³, K. Agatsuma¹⁴, N. Aggarwal¹⁵, O. D. Aguiar¹⁶, L. Aiello^{17,18}, A. Ain¹⁹, P. Ajith²⁰, B. Allen^{10,21,22}, G. Allen¹², A. Allocca^{23,25}, P. A. Altin²⁵, A. Amato²⁶, A. Ananyeva¹, S. B. Anderson¹, W. G. Anderson²¹, S. V. Angelova²⁷, S. Antier²⁸, S. Apperl¹, K. Arai¹, M. C. Araya¹, J. S. Areeda²⁹, N. Amaud^{28,30}, K. G. Arun³¹, S. Ascenzi^{32,33}, G. Ashton¹⁰, M. Ast³⁴, S. M. Aston⁷, P. Astone³⁵, D. V. Atallah³⁶, P. Aufmuth²², C. Aulbert¹⁰, K. AultONeal³⁷, C. Austin², A. Avila-Alvarez⁹, S. Babak³⁸, P. Bacon³⁹, M. K. M. Bader¹⁴, S. Bae⁴⁰, P. T. Baker⁴¹, F. Baldaccini^{52,43}, G. Ballardin³⁰, S. W. Ballmer⁴⁴, S. Banagiri⁴⁵, J. C. Barayoga¹, S. E. Barclay⁴⁶, B. C. Barish¹, D. Barker⁴⁷, K. Barket¹⁸, F. Barone^{3,4}, B. Barr⁴⁶, L. Barsotti¹⁵, M. Barsuglia³⁹, D. Barta⁴⁹, S. D. Barthelmy⁵⁰, J. Bartlett⁴⁷, I. Batros^{15,5}, R. Bassiri⁵², A. Basti^{32,44}, J. C. Batch⁴⁷, M. Bawaj^{53,43}, J. C. Bayley⁴⁶, M. Bazzaf^{4,55}, B. Bécsy⁵⁶, C. Beer¹⁰, M. Beijger⁵⁷, I. Belahcene²⁸, A. S. Bell⁴⁶, B. K. Berger¹, G. Bergmann¹⁰, J. J. Bero⁸, C. P. L. Berry⁵⁹, D. Bersanetti⁶⁰, A. Bertolini¹⁴, J. Betzwieser⁷, S. Biscas^{1,15}, S. Biscoveanu^{64,6}, A. Bisht²², M. Bitossi^{30,24}, C. Biwe⁴⁴, M. A. Bizouar²⁸, J. K. Blackburn¹, J. Blackman⁴⁸, C. D. Blair^{1,65}, D. G. Blain⁶⁵, R. M. Blair⁴⁷, S. Bloeme⁶⁶, O. Boec¹⁰, N. Bode¹⁰, M. Boer⁵⁷, G. Bogaert⁵⁷, A. Bohe³⁸, C. D. Blair^{1,65}, D. G. Blain⁶⁵, R. M. Blair⁴⁷, S. Bloeme⁶⁶, O. Boeck¹⁰, N. Bode¹⁰, M. Boer⁶⁷, G. Bogaert⁵⁷, A. Bohe³⁸, C. D. Blair^{1,65}, D. G. Blain⁶⁵, R. M. Blair⁴⁷, S. Bloeme⁶⁶, O. Boeck¹⁰, N. Bode¹⁰, M. Boer⁶⁷, G. Bogaert⁵⁷, A. Bohe³⁸, C. D. Blair^{1,65}, D. G. Blain⁶⁵, R. M. Blair

- 62 collaborations
 - (Can't fit in AstroPH author display)
- 3611 authors

D

S. L

М

- Author list is 10 pages long
 - (Normal ApJL total length is 4 pgs)
- 953 institutional affiliations
- Acknowledgements take 6 pgs
 - 4 authors are already dead

P. Gruning²⁸, G. M. Guidi^{121,122}, X. Guo⁸², A. Gupta⁶⁴, M. K. Gupta¹⁰⁵, K. E. Gushwa¹, E. K. Gustafson¹, R. Gustafson¹⁸, O. Halim^{18,17}, B. R. Hall⁶⁹, E. D. Hall¹⁵, E. Z. Hamilton³⁶, G. Hammond⁴⁶, M. Haney¹²³, M. M. Hanke¹⁰, J. Hanks⁴⁷, C. Hanna⁶⁴, M. D. Hanna³⁶, O. A. Hannuksela⁹³, J. Hanson⁷, T. Hardwick², J. Harms^{17,18}, G. M. Harry¹²⁴, I. W. Harry³⁸, M. J. Hart⁶⁶, C.-J. Haster⁹⁰, K. Haughian⁴⁶, J. Healy³⁸, A. Heidmann⁷¹, M. C. Heintze⁷, H. Heitmann⁶⁷, P. Hello²⁸, G. Hemming³⁰, M. Hendry⁴⁶, I. S. Heng⁶⁶, J. Hennig⁴⁶, A. W. Heptonstall¹, M. Heurs^{10,22}, S. Hild⁴⁶, T. Hinderet⁶⁶, D. Hoak³⁰, D. Hofman²⁶, K. Holt⁷, D. E. Holz⁹¹, P. Hopkins³⁶, C. Horst²¹, J. Hough⁴⁶, E. A. Houston⁴⁶, E. J. Howell⁶⁵, A. Hreibi⁶⁷, Y. M. Hu¹⁰, E. A. Huerta¹², D. Huet²⁸, B. Hughey³⁷, S. Husa¹⁰², S. H. Huutner⁴⁶, T. Huynh-Dinh⁷, N. Indik¹⁰, R. Inta⁸, G. Intin^{47,15}, H. N. Isa⁴⁶, J.-M. Isac¹¹, M. Isi¹, B. R. Iyer²⁰, K. Lizuni⁴⁷, T. Jacqmin⁷¹, K. Jani⁷⁷, P. Jaranowski¹²⁵, S. Jawahar⁶³, F. Jiménez-Forteza¹⁰², W. W. Johnson², D. I. Jones¹²⁶, R. Jones⁴⁶, R. J. G. Jonker¹⁴, L. Ju⁶⁵, J. Junker¹⁰, C. V. Kalaghatgi³⁶, V. Kalogera⁸⁹, B. Kamai¹, S. Kandhasamy⁷, G. Kang⁴⁰, J. B. Kanner¹, S. J. Kapadia²¹, S. Karki⁷⁰



PHYSICAL REVIEW D 90, 102002 (2014)

Multimessenger search for sources of gravitational waves and high-energy neutrinos: Initial results for LIGO-Virgo and IceCube

M. G. Aartsen,^{2,*} M. Ackermann,^{45,*} J. Adams,^{15,*} J. A. Aguilar,^{23,*} M. Ahlers,^{28,*} M. Ahrens,^{36,*} D. Altmann,^{22,*} T. Anderson,^{42,*} C. Arguelles,^{28,*} T. C. Arlen,^{42,*} J. Auffenberg,^{1,*} X. Bai,^{34,*} S. W. Barwick,^{25,*} V. Baum,^{29,*} J. J. Beatty,^{17,18,*} J. Becker Tjus,^{10,*} K.-H. Becker,^{44,*} S. BenZvi,^{28,*} P. Berghaus,^{45,*} D. Berley,^{16,*} E. Bernardini,^{45,*} A. Bernhard,^{31,*} D. Z. Besson,^{26,*} G. Binder,^{8,7,*} D. Bindig,^{44,*} M. Bissok,^{1,*} E. Blaufuss,^{16,*} J. Blumenthal,^{1,*} D. J. Boersma,^{43,*} C. Bohm,^{36,*} F. Bos,^{10,*} D. Bose,^{38,*} S. Böser,^{11,*} O. Botner,^{43,*} L. Brayeur,^{13,*} H.-P. Bretz,^{45,*} A. M. Brown,^{15,*} J. Casey,^{5,*} M. Casier,^{13,*} D. Chirkin,^{28,*} A. Christov,^{23,*} B. Christy,^{16,*} K. Clark,^{39,*} L. Classen,^{22,*} F. Clevermann,^{20,*} S. Coenders,^{31,*} D. F. Cowen,^{42,41,*} A. H. Cruz Silva,^{45,*} M. Danninger,^{36,*} J. Daughhetee,^{5,*} J. C. Davis,^{17,*} M. Day,^{28,*} J. P. A. M. de André,^{42,*} C. De Clercq,^{13,*} S. De Ridder,^{24,*} P. Desiati,^{28,*} K. D. de V. M. de With,^{9,*} T. DeYoung,^{42,*} J. C. Díaz-Vélez,^{28,*} M. Dunkman,^{42,*} R. Eagan,^{42,*} B. Eberhardt,^{29,*} B. Evented,^{28,*} S. Evented,^{28,*} S. Evented,^{42,*} R. Eagan,^{42,*} B. Eberhardt,^{29,*} B. Evented,^{28,*} S. Evented,^{28,*} S. Evented,^{42,*} R. Eagan,^{42,*} B. Eberhardt,^{29,*} B. Evented,^{28,*} S. Evented,^{28,*} S. Evented,^{42,*} R. Eagan,^{42,*} B. Evented,^{42,*} B. Evented,^{42,*} B. Evented,^{41,*} B. Evented,^{42,*} S. Ev M. de winn, ⁻¹. De young, ⁻¹. J. C. Diaz-Velez, ^{--.} M. Dunkman, ^{--.} R. Eagan, ^{--.} B. Eberhardt, ^{25.} B.
J. Eisch, ^{28.} S. Euler, ^{43.} P. A. Evenson, ^{32.} O. Fadiran, ^{28.} A. R. Fazely, ^{6.} A. Fedynitch, ^{10.} J. Feintzei T. Feusels, ^{24.} K. Filimonov, ^{7.} C. Finley, ^{36.} T. Fischer-Wasels, ^{44.} S. Flis, ^{36.} A. Franckowiak T. Fuchs, ^{20.} T. K. Gaisser, ^{32.} J. Gallagher, ^{27.} L. Gerhardt, ^{8.7.} D. Gier, ^{1.} L. Gladstonov, ^{11.} G. Goldschmidt, ^{8.*} G. Golup, ^{13.*} J. G. Gonzalez, ^{32.*} J. A. Goodman, ^{16.*} D. Góra, ^{45.*} D. Góra, ^{45.*} D. Gretskov, ^{1.*} J. C. Groh, ^{42.*} A. Groß, ^{31.*} C. Ha, ^{87.*} C. Haack, ^{1.*} A. Haj Ismetric F. Halzen, ^{28.*} K. Hanson, ^{12.*} D. Hebecker, ^{11.*} D. Heereman, ^{12.*} D. Heinen, ^{1.*} K. F. Halzen, ^{26,-} K. Hanson, ^{12,-} D. Hebecker, ^{11,*} D. Heereman, ^{12,*} D. Heinen, ¹S. Hickford, ^{15,*} G. C. Hill, ^{2,*} K. D. Hoffman, ^{16,*} R. Hoffmann, ^{44,*} A W. Huelsnitz, ^{16,*} P. O. Hulth, ^{36,*} K. Hultqvist, ^{36,*} S. Hussain ^{32,*} K. Jagielski, ^{1,*} G. S. Japaridze, ^{4,*} K. Jero, ^{28,*} O. Jlelati, ^{24,*} M. Kauer, ^{28,*} J. L. Kelley, ^{28,*} A. Kheirandiski, ^{16,*} R. Kohnen, ^{30,*} H. Kolanoski, ^{9,*} A. Koob, ^{1,*} L. Köpko, ^{20,*} G. Kohnen, ^{30,*} H. Kolanoski, ^{9,*} A. Koob, ^{1,*} L. Köpko, ^{20,*} D. T. Larsen, ^{12,*} M. J. Larson, ^{19,*} M. Lesi, J. Madsen, ^{35,*} G. Maggi, ^{13,*} R. Maruw, A. Meli, ^{24,*} T. Moures, ^{12,*} S. Missing, ^{13,*} R. Maruw, ^{12,*} M. Moltan, ^{24,*} T. Moures, ^{12,*} S. Missing, ^{14,*} T. Moures, ^{15,*} S. Missing, ^{15,*} S. Hussing, ^{15,} . Medici, 19, L. Mohrmann,⁴ T. Montaruli,^{23,*} R. Morse,²⁸ A1,^{21,*} D. R. Nygren,^{8,*} alczewski,^{40,*} L. Paul,^{1,*} E. Pinat,^{12,*} J. Posselt,^{44,*} A. Obertacke,44,* S Ö. Penek,1,* neez,^{23,*} K. Rawlins,^{3,*} P. Redl,^{16,*} P. B. Price I. Rees 2¹
I. Rees 2²
I. Rees 2²< del,^{28,*} S. Robertson,^{2,*} J. P. Rodrigues,^{28,*} I. Rees 25 W. G. Anderson, ^{60,†} K. Arai, ^{70,†} M. C. Araya, ^{70,†} C. Arceneaux, ^{60,†} J. S. Areeda, ^{60,†} P. T. Baker, ^{70,†} G. Ballardin, ^{71,†} P. Astone, ^{71,†}
P. Aufmuth, ^{61,†} H. Augustus, ^{68,†} C. Aulbert, ^{54,†} B. E. Aylott, ^{68,†} S. Babak, ^{60,†} P. T. Baker, ^{70,†} G. Ballardin, ^{71,†}
S. W. Ballmer, ^{59,†} J. C. Barayoga, ^{46,†} M. Barbet, ^{50,†} B. C. Barish, ^{46,†} D. Barker, ^{72,†} F. Barone, ^{48,49,†} B. Barr, ^{73,†}
L. Barsotti, ^{56,†} M. Barsuglia, ^{74,†} M. A. Barton, ^{72,†} I. Bartos, ^{75,†} R. Bassini, ^{64,†} A. Basti, ^{76,63,†} J. C. Batch, ^{72,†}
J. Bauchrowitz, ^{54,†} Th. S. Bauer, ^{55,†} C. Baune, ^{54,†} V. Bavigadda, ^{71,†} B. Behnke, ^{69,†} M. Bejger, ^{77,†} M. G. Beker, ^{55,†}
C. Belczynski, ^{78,†} A. S. Bell, ^{73,†} C. Bell, ^{73,†} G. Bergmann, ^{54,†} D. Bersanetti, ^{79,80,†} A. Bertolini, ^{55,†} J. Betzwieser, ^{51,†}
I. A. Bilenko, ^{81,†} G. Billingsley, ^{46,†} J. Birch, ^{51,†} S. Biscans, ^{56,†} M. Bitossi, ^{63,†} C. Biwer, ^{59,†} M. A. Bizouard, ^{82,‡} E. Black, ^{46,†}
J. K. Blackburn, ^{46,†} L. Blackburn, ^{83,†} D. Blair, ^{84,†} S. Bloemen, ^{55,85,†} O. Bock, ^{54,†} T. P. Bodiya, ^{56,†} M. Boer, ^{86,†}



PHYSICAL REVIEW D 90, 102002 (2014)

Multimessenger search for sources of gravitational waves and high-energy neutrinos: Initial results for LIGO-Virgo and IceCube

M. G. Aartsen,^{2,*} M. Ackermann,^{45,*} J. Adams,^{15,*} J. A. Aguilar,^{23,*} M. Ahlers,^{28,*} M. Ahrens,^{36,*} D. Altmann,^{22,*} T. Anderson,^{42,*} C. Arguelles,^{28,*} T. C. Arlen,^{42,*} J. Auffenberg,^{1,*} X. Bai,^{34,*} S. W. Barwick,^{25,*} V. Baum,^{29,*} J. J. Beatty,^{17,18,*} J. Becker Tjus,^{10,*} K.-H. Becker,^{44,*} S. BenZvi,^{28,*} P. Berghaus,^{45,*} D. Berley,^{16,*} E. Bernardini,^{45,*} A. Bernhard,^{31,*} D. Z. Besson,^{26,*} G. Binder,^{8,7,*} D. Bindig,^{44,*} M. Bissok,^{1,*} E. Blaufuss,^{16,*} J. Blumenthal,^{1,*} D. J. Boersma,^{43,*} C. Bohm,^{36,*} F. Bos,^{10,*} D. Bose,^{38,*} S. Böser,^{11,*} O. Botner,^{43,*} L. Brayeur,^{13,*} H.-P. Bretz,^{45,*} A. M. Brown,^{15,*} J. Casey,^{5,*} M. Casier,^{13,*} D. Chirkin,^{28,*} A. Christov,^{23,*} B. Christy,^{16,*} K. Clark,^{39,*} L. Classen,^{22,*} F. Clevermann,^{20,*} S. Coenders,^{31,*} D. F. Cowen,^{42,41,*} A. H. Cruz Silva,^{45,*} M. Danninger,^{36,*} J. Daughhetee,^{5,*} J. C. Davis,^{17,*} M. Day,^{28,*} J. P. A. M. de André,^{42,*} C. De Clercq,^{13,*} S. De Ridder,^{24,*} P. Desiati,^{28,*} K. D. de V. M. de With,^{9,*} T. DeYoung,^{42,*} J. C. Díaz-Vélez,^{28,*} M. Dunkman,^{42,*} R. Eagan,^{42,*} B. Eberhardt,^{29,*} B. Evitation of the strange A Market Market

J. Madsen, ⁵⁵, G. Maggi ^{11,5}, R. Maruk, A. Meliz, ²⁴, T. Meures, ⁷², S. Mi T. Montaruli, ²³, R. Morse A. Obertacke, ⁴⁴, S O. Penek, ^{1,4}, J. P. R. Maruk, P. B. Price I. Rees ²⁸ I. Rees ²⁹ I. Rees ²⁰ I. Rees W. G. Anderson, ^{60,†} K. Arai, ^{76,†} M. C. Araya, ^{46,†} C. Arceneaux, ^{60,†} J. S. Areeda, ^{60,†} S. Ast, ^{61,†} S. M. Aston, ^{51,†} P. Astone, ^{61,†} P. Aufmuth, ^{61,†} H. Augustus, ^{68,†} C. Aulbert, ^{54,†} B. E. Aylott, ^{68,†} S. Babak, ^{69,†} P. T. Baker, ^{70,†} G. Ballardin, ^{71,†} S. W. Ballmer, ^{59,†} J. C. Barayoga, ^{46,†} M. Barbet, ^{50,†} B. C. Barish, ^{46,†} D. Barker, ^{72,†} F. Barone, ^{48,49,†} B. Barr, ^{73,†} L. Barsotti, ^{56,†} M. Barsuglia, ^{74,†} M. A. Barton, ^{72,†} I. Bartos, ^{75,†} R. Bassiri, ^{64,†} A. Basti, ^{76,63,†} J. C. Batch, ^{72,†} J. Bauchrowitz, ^{54,†} Th. S. Bauer, ^{55,†} C. Baune, ^{54,†} V. Bavigadda, ^{71,†} B. Behnke, ^{69,†} M. Bejger, ^{77,†} M. G. Beker, ^{55,†} C. Belczynski, ^{78,†} A. S. Bell, ^{73,†} C. Bell, ^{73,†} G. Bergmann, ^{54,†} D. Bersanetti, ^{79,80,†} A. Bertolini, ^{55,†} J. Betzwieser, ^{51,†}
I. A. Bilenko, ^{81,†} G. Billingsley, ^{46,†} J. Birch, ^{51,†} S. Biscans, ^{56,†} M. Bitossi, ^{63,‡} C. Biwer, ^{59,†} M. A. Bizouard, ^{82,‡} E. Black, ^{46,†} J. K. Blackburn, ^{46,†} L. Blackburn, ^{84,4,†} S. Bloemen, ^{55,85,†} O. Bock, ^{54,†} T. P. Bodiya, ^{56,†} M. Boer, ^{86,†}

